

1. Consider the evolution operator $\hat{U}(t, t_0)$ of a quantum particle. The coordinate-basis matrix of this operator

$$U(\mathbf{x}_1, t_1; \mathbf{x}_0, t_0) = \langle \mathbf{x}_1 | \hat{U}(t_1, t_0) | \mathbf{x}_0 \rangle \quad (1)$$

is called the *evolution kernel* or the *propagation amplitude* from \mathbf{x}_0 to \mathbf{x}_1 in time $t_1 - t_0$. This kernel describes the time evolution of the particle's wave-function as

$$\psi(\mathbf{x}_1, t_1) = \int d^3\mathbf{x}_0 U(\mathbf{x}_1, t_1; \mathbf{x}_0, t_0) \times \psi(\mathbf{x}_0, t_0). \quad (2)$$

Now consider a free non-relativistic spinless particle with Hamiltonian

$$\hat{H} = \frac{\hat{\mathbf{p}}^2}{2M}. \quad (3)$$

- (a) Show that the evolution kernel for this particle is

$$U(\mathbf{x}_1, t_1; \mathbf{x}_0, t_0) = \left(\frac{M}{2\pi i \hbar (t_1 - t_0)} \right)^{3/2} \times \exp \left(\frac{i}{\hbar} \frac{m(\mathbf{x}_1 - \mathbf{x}_0)^2}{2(t_1 - t_0)} \right). \quad (4)$$

- (b) Spell out the unitarity conditions $\hat{U}^\dagger \hat{U} = \hat{U} \hat{U}^\dagger = 1$ in terms of the evolution kernel (1), specifically in terms of integrals of the form $\int d^3\mathbf{x} U^*(\dots) U(\dots)$.
- (c) Now verify these conditions by explicit integration.

2. A classical charged particle in a magnetic field has *canonical momentum*

$$\mathbf{p} = m\mathbf{v} + \frac{Q}{c} \mathbf{A}(\mathbf{x}_{\text{particle}}) \quad (5)$$

which is quite different from the usual *kinematic momentum* $\vec{\pi} = m\mathbf{v}$, and its classical Hamiltonian is

$$H(\mathbf{x}, \mathbf{p}) = \frac{\vec{\pi}^2}{2m} = \frac{1}{2m} \left(\mathbf{p} - \frac{Q}{c} \mathbf{A}(\mathbf{x}) \right)^2. \quad (6)$$

In quantum mechanics, it's the canonical momentum operators \hat{p}_i ($i = x, y, z$) which obeys

the usual commutation relations

$$[\hat{x}_i, \hat{x}_j] = 0, \quad [\hat{p}_i, \hat{p}_j] = 0, \quad [\hat{x}_i, \hat{p}_j] = i\hbar\delta_{ij} \quad (7)$$

so in the coordinate basis they act as $\hat{p}_i\psi(\mathbf{x}) = -i\hbar(\partial/\partial x_i)\psi(\mathbf{x})$. On the other hand, the kinematic momenta

$$\hat{\pi}_i \stackrel{\text{def}}{=} \hat{p}_i - \frac{Q}{c} A_i(\hat{x}, \hat{y}, \hat{z}) \quad (8)$$

act in a more complicated fashion and obey more complicated commutation relations

$$[\hat{x}_i, \hat{x}_j] = 0, \quad [\hat{x}_i, \hat{\pi}_j] = i\hbar\delta_{ij}, \quad [\hat{\pi}_i, \hat{\pi}_j] = \frac{i\hbar Q}{c} \epsilon_{ijk} B_k(\hat{x}, \hat{y}, \hat{z}). \quad (9)$$

Finally, the Hamiltonian operator \hat{H} follows from the classical Hamiltonian (6) as

$$\hat{H} = \frac{\vec{\hat{\pi}}^2}{2m} = \frac{1}{2m} (\hat{\mathbf{P}} - \mathbf{A}(\hat{\mathbf{x}}))^2. \quad (10)$$

- (a) Unless you have attended my extra lecture on September 9, read the parts of [my notes on classical mechanics and canonical quantization](#) where I explain eqs. (5) through (10). Specifically, pages 5–7 where I explain the classical mechanics of a charged particle, and pages 8–9 where I explain the commutation relations.
- (b) Use the commutation relations (9) to derive the Ehrenfest equations for the quantum charged particle. Specifically, show that

$$\frac{d}{dt} \langle \hat{\mathbf{x}} \rangle = \frac{1}{m} \langle \vec{\hat{\pi}} \rangle \quad \text{and} \quad \frac{d}{dt} \langle \vec{\hat{\pi}} \rangle = \frac{Q}{2mc} \langle \vec{\hat{\pi}} \times \hat{\mathbf{B}} - \hat{\mathbf{B}} \times \vec{\hat{\pi}} \rangle \quad (11)$$

where $\hat{\mathbf{B}} \stackrel{\text{def}}{=} \mathbf{B}(\hat{x}, \hat{y}, \hat{z})$.

- (c) Now let's subject the particle to both electric and magnetic fields and allow both fields to be time-dependent, thus time-dependent Hamiltonian

$$\hat{H}(t) = Q\Phi(\hat{\mathbf{x}}, t) + \frac{1}{2m}(\hat{\mathbf{p}} - \mathbf{A}(\hat{\mathbf{x}}, t))^2 \quad (12)$$

Show that in this case, the Ehrenfest equations become

$$\frac{d}{dt}\langle\hat{\mathbf{x}}\rangle = \frac{1}{m}\langle\vec{\hat{\pi}}\rangle \quad \text{and} \quad \frac{d}{dt}\langle\vec{\hat{\pi}}\rangle = Q\langle\vec{\hat{\mathbf{E}}}\rangle + \frac{Q}{2mc}\langle\vec{\hat{\pi}}\times\hat{\mathbf{B}} - \hat{\mathbf{B}}\times\vec{\hat{\pi}}\rangle. \quad (13)$$

Hint: use Heisenberg–Dirac equations for the time-dependent operators, and remember that in a time-dependent vector potential $\mathbf{A}(\mathbf{x}, t)$, the kinematic momentum operator (8) becomes *explicitly* time-dependent,

$$\frac{\partial}{\partial t}\vec{\hat{\pi}} = -\frac{Q}{c}\frac{\partial\mathbf{A}}{\partial t}(\hat{\mathbf{x}}, t). \quad (14)$$

3. Finally, a refresher of undergraduate-level theory of orbital angular momentum

$$\hat{\mathbf{L}} \stackrel{\text{def}}{=} \hat{\mathbf{x}} \times \hat{\mathbf{p}}, \quad \text{i. e.,} \quad \hat{L}_i \stackrel{\text{def}}{=} \epsilon_{ijk}\hat{x}_j\hat{p}_k. \quad (15)$$

It is also a good drill for the use of the canonical commutation relations

$$[\hat{x}_i, \hat{x}_j] = 0, \quad [\hat{p}_i, \hat{p}_j] = 0, \quad [\hat{x}_i, \hat{p}_j] = i\hbar\delta_{ij}. \quad (16)$$

Without using coordinate-basis or any other wave functions, use the relations (16) to show that:

- (a) $[\hat{x}_i, \hat{L}_j] = i\hbar\epsilon_{ijk}\hat{x}_k$;
- (b) $[\hat{p}_i, \hat{L}_j] = i\hbar\epsilon_{ijk}\hat{p}_k$;
- (c) $[\hat{L}_i, \hat{L}_j] = i\hbar\epsilon_{ijk}\hat{L}_k$ and therefore $\hat{\mathbf{L}} \times \hat{\mathbf{L}} = i\hbar\hat{\mathbf{L}}$;
- (d) $[\hat{\mathbf{p}}^2, \hat{\mathbf{L}}] = 0 = [f(\hat{r}), \hat{\mathbf{L}}]$ for any function f of the radius $\hat{r} = (\hat{\mathbf{x}}^2)^{1/2}$;

(e) $[\hat{\mathbf{L}}, \hat{\mathbf{L}}^2] = 0$;

(f) $\hat{\mathbf{p}}^2 = \hat{p}_r^2 + \hat{r}^{-2} \hat{\mathbf{L}}^2$, where $\hat{p}_r \stackrel{\text{def}}{=} \frac{1}{2} \left\{ \frac{\hat{x}_i}{r}, \hat{p}_i \right\}$ (note: \hat{p}_r so defined is hermitian).

Now, let us make use of the coordinate-basis wave functions.

(g) Given a wave-function $\Psi(r, \theta, \phi) = \langle r, \theta, \phi | \Psi \rangle$ in the spherical-coordinate basis, compute $\langle r, \theta, \phi | \hat{p}_r | \Psi \rangle$ and $\langle r, \theta, \phi | \hat{p}_r^2 | \Psi \rangle$.

(h) Finally, use the above results to calculate $\langle r, \theta, \phi | \hat{\mathbf{L}}^2 | \Psi \rangle$.

Hint: In the coordinate basis $\hat{\mathbf{p}}^2 = -\hbar^2 \nabla^2$; look up in any E&M textbook how the Laplacian ∇^2 acts in spherical coordinates, then use (f) and (g).