

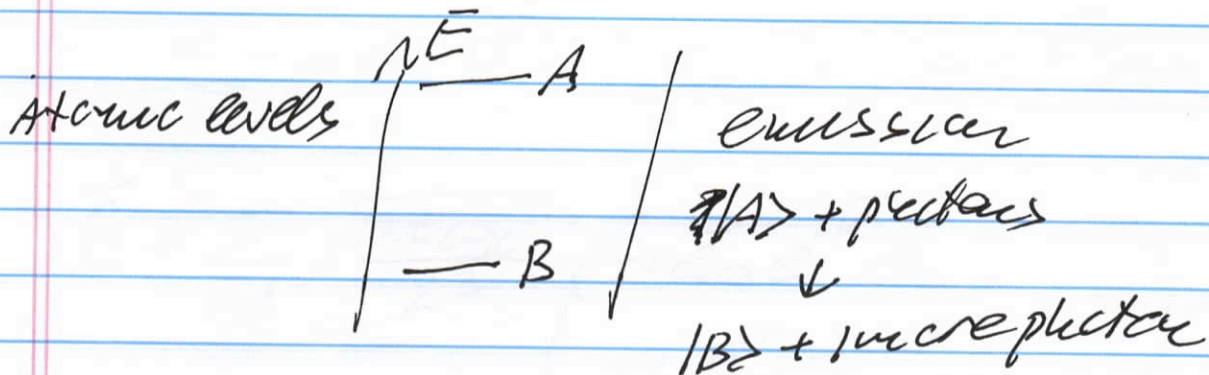
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# EM fields + Atoms (S)

$$\hat{H} = \hat{H}_{EM} + \hat{H}_{atom} + \hat{H}_{int.}$$

$\hat{H}_{int}$  causes transitions:

emission & absorption of light



absorption

$|B\rangle + \text{photons} \rightarrow |A\rangle + \text{less photons.}$

## Fermi Golden rule

transition rate  $\Gamma = \frac{d(\text{probability})}{d(\text{time})}$

$$\Gamma = \frac{2\pi}{\hbar} |\langle \text{final} | \hat{H}_{int} | \text{initial} \rangle|^2 \rho$$

$\rho =$  density of final states

$$= \frac{d \# \text{ final states}}{d E_{\text{final}}}$$

For atom size  $\ll$  wavelength

$$\hat{H}_{int} = -\hat{\vec{d}} \cdot \hat{\vec{E}}(\vec{r}_{atom})$$

atom is electric dipole.

$$\langle f_{em} | \hat{H}_{int} | u \rangle = \langle f_{em} |$$

$$= -2 \langle f_{em} | \hat{\vec{d}} | u \rangle_{atom} \cdot \langle E_{f_{em}} | \hat{\vec{E}} | E_{u_{em}} \rangle$$

$$\hat{\vec{E}} = \sum_{k\lambda} \sqrt{\frac{2\pi\hbar\omega_k}{L^3}} \begin{pmatrix} -i\vec{e}_{k\lambda} e^{i\vec{k}\cdot\vec{r}} a_{k\lambda} \\ +i\vec{e}_{k\lambda}^* e^{-i\vec{k}\cdot\vec{r}} a_{k\lambda}^\dagger \end{pmatrix}$$

absorption:  $\langle \{u\} - 1_{u_x} | \hat{\vec{E}} | \{u\} \rangle$

$$= -i \sqrt{\frac{2\pi\hbar\omega_k}{L^3}} (-i e^{i\vec{k}\cdot\vec{r}} \vec{e}_{k\lambda}) \cdot \sqrt{u_{k\lambda}}$$

emission  $\langle \{u_{k\lambda}\} + 1_{u_x} | \hat{\vec{E}} | \{u\} \rangle =$

$$= \sqrt{\frac{2\pi\hbar\omega_k}{L^3}} (i e^{-i\vec{k}\cdot\vec{r}} \vec{e}_{k\lambda}) \cdot \sqrt{u_{k\lambda} + 1}$$

atom:  $\langle B | \hat{\vec{d}} | A \rangle = \langle A | \hat{\vec{d}} | B \rangle^*$

$$\Gamma_{emission} = \Gamma_0 \times (u_{k\lambda} + 1) (u_{k\lambda} + 1)$$

$$\Gamma_{absorption} = \Gamma_0 \times u_{k\lambda}$$

same  $\Gamma_0$

$$\Gamma_{\text{absorb}} = \Gamma_0 \times n_{\text{up}}$$

$$\Gamma_{\text{emiss}} = \Gamma_0 \times (n_{\text{up}} + 1)$$

$$= \Gamma_0 + \Gamma_0 \times n_{\text{up}}$$

←  
spontaneous  
emission

↓  
stimulated  
emission.

Einstein (1917)

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Einstein is assuring:

detailed balance in thermal  
equilibrium

photons & atoms in a cavity.

net transition rates between atoms

in excite state A  
and less excited state B

should be equal in both directions

$$\begin{aligned} \# \text{ Atoms } (A) \times \Gamma_{\text{emiss}} &= \\ &= \# \text{ Atoms } (B) \times \Gamma_{\text{absorb}} \end{aligned}$$

$$\begin{aligned} \frac{\Gamma_{\text{emiss}}}{\Gamma_{\text{absorb}}} &= \frac{\# \text{ Atoms } (B)}{\# \text{ Atoms } (A)} = e^{-\frac{(E_B - E_A)}{k_B T}} \\ &= e^{-\frac{\hbar \omega}{k_B T}} \end{aligned}$$

$$\frac{\Gamma_{\text{em}}}{\Gamma_{\text{abs}}} = e^{+\hbar\omega/k_B T}$$

by Planck law  $\langle n_{\gamma} \rangle = \frac{1}{e^{\hbar\omega/k_B T} - 1}$

$$\Rightarrow \frac{\langle n_{\gamma} \rangle + 1}{\langle n_{\gamma} \rangle} = e^{\hbar\omega/k_B T}$$

$$\frac{\Gamma_{\text{emission}}}{\Gamma_{\text{absorption}}} = \frac{\langle n_{\gamma} \rangle + 1}{\langle n_{\gamma} \rangle}$$

$$\Gamma_{\text{abs}} = \langle n_{\gamma} \rangle \cdot \Gamma_0 \rightarrow n_{\gamma} \Gamma_0$$

$$\Downarrow$$
$$\Gamma_{\text{emission}} = (\langle n_{\gamma} \rangle + 1) \Gamma_0 \rightarrow (n_{\gamma} + 1) \Gamma_0$$

non-equilibrium: # atoms (A) > # atoms (B)

→ population inversion

→ stimulated emission can amplify light.

A particle in 1 dimension

$$\hat{H} = \frac{\hat{p}^2}{2m} + V(x).$$

Schrödinger eq -  $\hat{H}|\psi\rangle = E|\psi\rangle$

$$\hookrightarrow -\frac{\hbar^2}{2m} \psi''(x) + V(x)\psi(x) = E\psi(x).$$

$$\left( \frac{\psi''(x)}{\psi(x)} = \frac{2m}{\hbar^2} (V(x) - E) \right)$$

$\hookrightarrow$  2<sup>nd</sup> order diff. eq - n

$\rightarrow$  2 independent solutions.

valid or invalid states  $|\psi\rangle$

depending on asymptotic

behavior @  $x \rightarrow \pm\infty$ .

$x \rightarrow +\infty$

if  $V(+\infty) > E$

asymptotically  $\psi(x) = A e^{\alpha x} + B e^{-\alpha x}$

$$\alpha = \frac{\sqrt{2m(V-E)}}{\hbar}$$

$e^{-\alpha x}$  is good.  $\rightarrow$  normalizable  $\psi(x)$ .

$e^{+\alpha x}$  is bad: it blows up @  $\infty$ .

$\rightarrow$  unnormalizable  $\psi(x)$

but it's not even asymptotic

General rule: solutions  
which blow up exponentially  
(or faster) @  $x \rightarrow \pm\infty$  are ~~not~~ good.

(f)  $V(+\infty) < E$

Asymptotic  $\psi(x) = A e^{cx} + B e^{-cx}$

$\rightarrow$  2 independent unnormalizable  
solutions, both are ok.

ditto  $x \rightarrow -\infty$

- (1)  $E < \min(V(+\infty), V(-\infty))$
- (2)  $E$  between  $V(+\infty)$  &  $V(-\infty)$
- (3)  $E > \max(V(+\infty), V(-\infty))$

(1): Each end  $x \rightarrow \pm\infty$  selects  
a particular linear combination  
of 2 solutions  $\psi_1(x)$  &  $\psi_2(x)$   
of the Schrödinger eqn.

for general  $E$ , the 2 ends select different  
combinations of  $\psi_1$  &  $\psi_2$   
 $\rightarrow$  no good solutions.

for some discrete  $E$ , both ends select  
same solution.

(1) For  $E < \min(V(+\infty), V(-\infty))$

The energy spectrum is discrete  
& non-degenerate.

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(2)  $V(-\infty) < E < V(+\infty)$

"Boundary" condition @  $x \rightarrow -\infty$

select one good solution.

(1 particular case of  $\psi_1, \psi_2$ )

OK, both solutions are OK @  $x \rightarrow +\infty$

$\rightarrow \exists$  <sup>good</sup> 1 solution  $\forall E$ .

$\rightarrow$  continuous, non-degenerate  
spectrum.

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(2')  $V(-\infty) > E > V(+\infty)$

$\rightarrow$  same story.

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(3)  $E > \text{both } V(+\infty), V(-\infty)$

$\rightarrow$  either solution  $\psi_1$  or  $\psi_2$

is OK @ both ends.

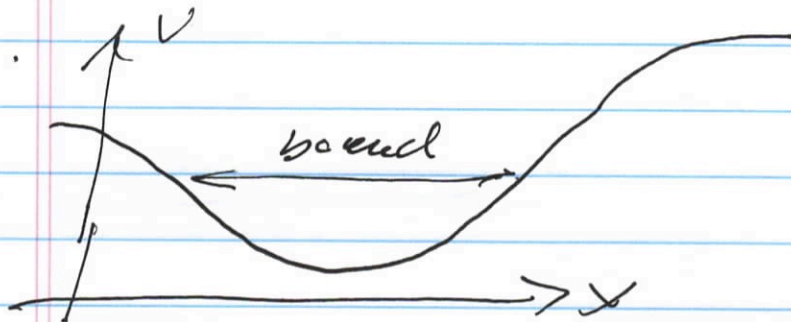
$\rightarrow$  continuous, doubly degenerate  
spectrum.

# Discretized Spectrum

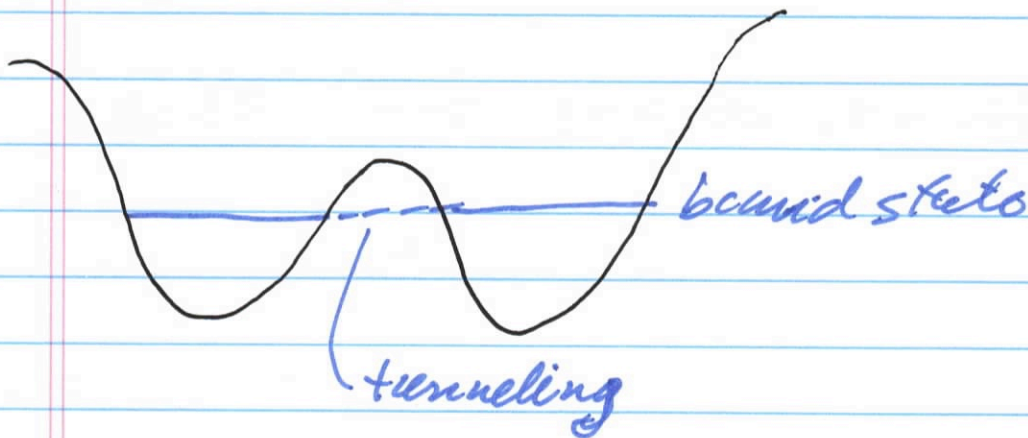
→ bound states

# Continuous spectrum

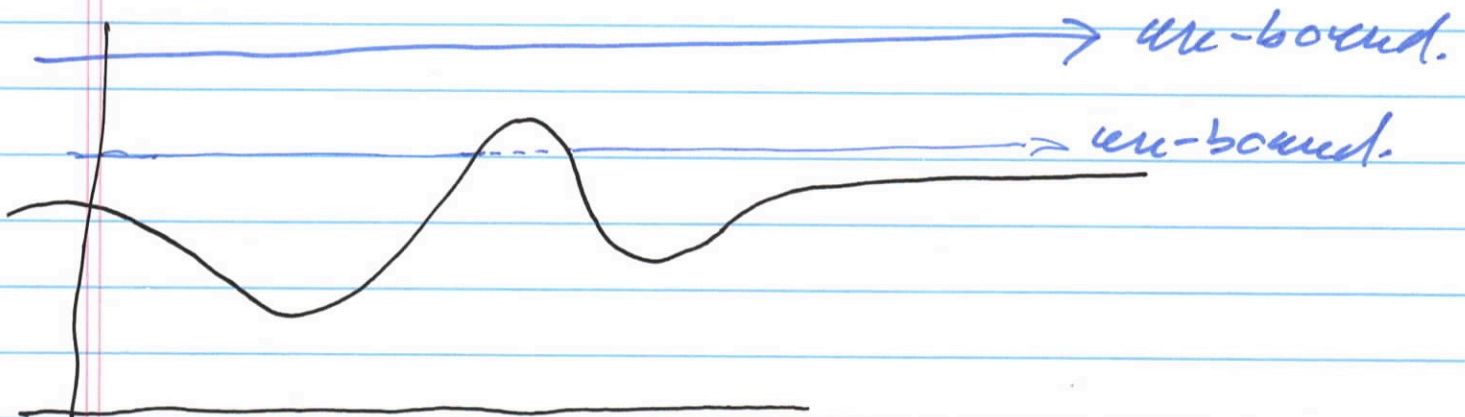
→ unbound states



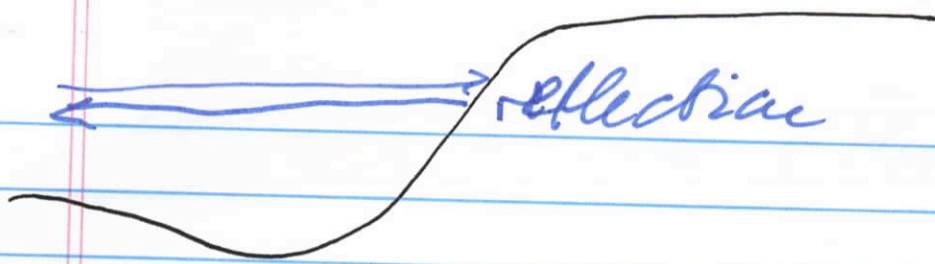
@  $x$ :  $\psi$  decreases exponentially beyond the classical bounds.



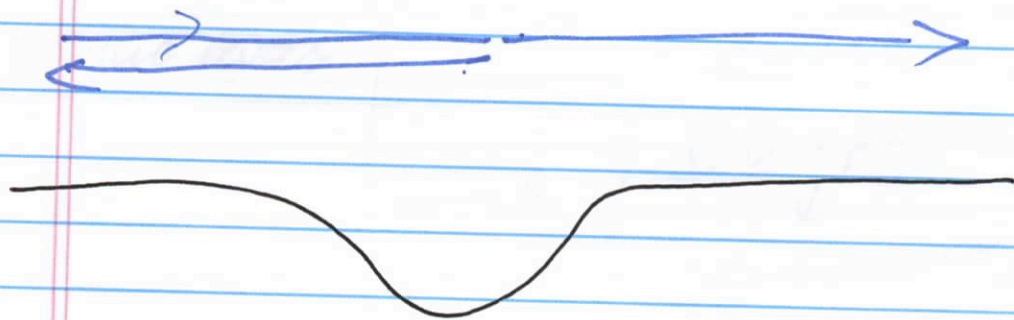
unbound







~~was~~ motion unbound @ left end.



both transmission & reflection